# ON THE STABILITY OF ROTATION OF A RIGID BODY SUSPENDED ON A STRING AND POSSESSING an ELLIPSOIDAL CAVITY COMPLETELY FILLED WITH AN IDEAL INCOMPRESSIBLE FLUID 

# (OR USTOICHIVOSTI VRASHCHENIIA NA STRUNE TVERDOGO TELA S ELLIPSOIDAL'NOI POLOST'IU, TSELIKOM NAPOLNENNOI IDEAL'NOI NESZHIMAEMOI ZHIDKOST'IU) 

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The present study of the stability of the rotation on a string of a solid with a cavity filled with an ideal fluid, again uses the method developed in [1]. The principle of this method consists in the description of the motion of the flaid with respect to a system of coordinates fixed with respect to the rigid body, performing motion under the action of the fluid, its own weight and reaction forces of the suspension.

The problem considered below was, with the flaid absent, investigated in [2]. Another limiting case corresponding to a string of zero length is the well known problem of Sobolev, investigated in [1 and 3]. We shall utilise some of the methods from the above works.

1. Let a symmetric rigid body suspended on an ideally flexible, inextensible and inertialess string of length $i$, rotate with a constant angular velocity $\omega$ about a vertical axis in a steady motion (Fig. 1).Inside the body, there is a cavity, which has the form of an ellipsoid of revolation, and which is completely filled with an ideal incompressible fluid. The axes of symmetry of the cavity and of the body coincide. During the steady motion the fluid rotates with the body, as if they were a single rigid body. We shall investigate the stability of such a stationary motion.
2. We shall consider first the derivation of the differential equations of the body in question, assuming that this motion does not differ much from the above mentioned stationary rotation whose angular velocity about a vertical axis is $\omega$.

We shall introduce the fixed system of coordinates $\xi \eta \zeta$ which has the vertical axis $\zeta$


FIG. 1
and its origin $O_{2}$ at the point of attachment of the string to a fixed support, and also the moving system $\xi^{\circ} \eta^{\circ} \zeta^{\circ}$ with its origin $O_{2}$ at the point at which the body is tied to the string (Fig. 2).

The corresponding axes of these systems are parallel. At the point $O$ we shall also locate the origin of the system of coordinates $x y z$ rigidly connected to the body. Axis $z$ of this system will coincide with the axes of symme try of the body and the cavity, and the axes $x$ and $y$ will lie in the plane perpendicular to the axis $z$ so as to form together a rectangular system of coordinates.

The position of the string with respect to the system $\boldsymbol{\xi} \eta \boldsymbol{\zeta}$ will be defined by two angles $\lambda$ and $\mu$; at the same time, $\lambda$ is the angle between the axis $\zeta$ and the projection on the axis $\eta \zeta$ of the straight line directed upwards along the string, while $\mu$ is the angle between that line and the plane $\eta \zeta$ (Fig. 2).

The position of the solid body with respect to the system $\xi^{\circ} \eta^{\circ} \zeta^{\circ}$ is determined by three Enler-Krylov angles (Fig. 3):


FIG. 2


FIG. 3
the angle $a$ between the axis $\zeta^{\circ}$ and the projection of the axis of symmetry of the body (axis z) on the plane $\eta^{0} \zeta^{\circ}$; the angle $\beta$ between the axis $z$ and the same plane $\eta^{\circ} \zeta^{\circ}$ and the angle $\varphi$ between the axis $x$ and the auxiliary axis $\xi^{\prime \prime}$ which is at the right anglea to the $z$-axia and lies in the plane $\xi^{\circ} z$.

The cosine table of the angles between the systems of coordinates $\xi \eta \zeta$ and $x y z$ or (which is the same), between the $\xi^{\circ} \eta^{\circ} \zeta^{\circ}$ and $x y z$ in the above case have the form

$$
\begin{array}{cccc}
\xi & n & \zeta \\
x & \cos \beta \cos \varphi & \sin \alpha \sin \beta \cos \varphi+ & -\sin \beta \cos \alpha \cos \varphi+ \\
& +\cos \alpha \sin \varphi & +\sin \alpha \sin \varphi \tag{2.1}
\end{array}
$$

| $y$ | $-\cos \beta \sin \varphi$ | $-\sin \alpha \sin \beta \sin \varphi+$ | $+\sin \beta \cos \alpha \sin \varphi+$ |
| :---: | :---: | :---: | :---: |
| $z$ | $\sin \beta$ | $-\sin \alpha \cos \beta$ | $+\sin \alpha \cos \varphi$ |
| $z$ | $-\cos \alpha \cos \beta$ |  |  |

It is easy to see from Fig. 3 that the projections $\omega_{x}, \omega_{y}$ and $\omega_{z}$ of the angular velocity of the system of coordinates $x y z$ on its own axes will be

$$
\begin{gather*}
\omega_{x}=\frac{d \alpha}{d t} \cos \beta \cos \varphi+\frac{d \beta}{d t} \sin \varphi \\
\omega_{y}=-\frac{d \alpha}{d t} \cos \beta \sin \varphi+\frac{d \beta}{d t} \cos \varphi  \tag{2.2}\\
\omega_{z}=\frac{d \alpha}{d t} \sin \beta+\frac{d \varphi}{d t}
\end{gather*}
$$

For future use we shall also define the coordinates of the point $O$ at which the body is attached to the string in terms of the aystem $\xi \eta \zeta$. We have

$$
\begin{equation*}
\xi_{0}=-l \sin \mu, \quad \eta_{0}=l \cos \mu \sin \lambda, \quad \xi_{0}=-l \cos \mu \cos \lambda \tag{2.3}
\end{equation*}
$$

The expression for the kinetic energy of the solid can be written in the form

$$
T=\frac{1}{2} m v_{0}^{2}+m\left|\begin{array}{lll}
\omega_{x} & \omega_{y} & \omega_{z} \\
x_{G} & y_{G} & z_{G} \\
\left(v_{0}\right)_{x} & \left(v_{0}\right)_{y} & \left(v_{0}\right)_{z}
\end{array}\right|+\frac{1}{2}\left[\left(A+m l_{1}^{2}\right)\left(\omega_{x}^{2}+\omega_{y}^{2}\right)+C \omega_{z}^{2}\right]
$$

Here $m$ is the mass of the rigid body; $A-m l_{1}^{2}=B+m l_{1}^{2}$ and $C$ are its moments of inertia with respect to the axes $x, y$ and $z ; v_{0}$ is the absolute velocity of the origin $O$ of the coordinates; $x_{G}, y_{G}, z_{G}$ are the coordinates of the center of gravity of the body in the aystem $x y z ; l_{1}$ is the distance between the center of gravity of the body and the point at which it is attached to the string (Fig. 1).

From the expression (2.4) and the formulas (2.2), and (2.3), we obtain an explicit expression for the kinetic energy interms of generalized coordinates $\alpha, \beta, \lambda$, and $\mu$ and generalized velocities $d \alpha / d t, d \beta / d t, d \lambda / d t, d \mu / d t$, and $d \varphi / d t$ (the generalized coordinate $\varphi$ is cyclic), which is

$$
\begin{gather*}
T=\frac{1}{2} m\left[l^{2}\left(\frac{d \mu}{d t}\right)^{2}+l^{2} \cos ^{2} \mu\left(\frac{d \lambda}{d t}\right)^{2}\right]+m l_{1} l \frac{d \alpha}{d t} \cos \beta\left[\frac{d \mu}{d t} \sin \mu \sin (\alpha-\lambda)+\right. \\
\left.+\frac{d \lambda}{d t} \cos \mu \cos (\alpha-\lambda)\right]+m l_{1} l \frac{d \beta}{d t}\left[\frac{d \mu}{d t} \cos \beta \cos \mu+\right. \\
\left.+\frac{d \mu}{d t} \sin \mu \sin \beta \cos (\alpha-\lambda)-\frac{d \lambda}{d t} \cos \mu \sin \beta \sin (\alpha-\lambda)\right]+  \tag{2.5}\\
+ \\
+\frac{1}{2}\left(A+m l_{1}^{2}\right)\left[\left(\frac{d \alpha}{d t}\right)^{2} \cos ^{2} \beta+\left(\frac{d \beta}{d t}\right)^{2}\right]+\frac{1}{2} C\left[\frac{d x}{d t} \sin \beta+\frac{d \varphi}{d t}\right]^{2}
\end{gather*}
$$

We shall now express the equations of motion of the solid by using the second Lagrange's method. Restricting ourselves to the infinitesimals of the first order with respect to the coordinates $\lambda, \mu, \alpha$, and $\beta$ and their time derivatives, we obtain the following set of linear differential equations with constant coefficients:

$$
\begin{gather*}
m l\left(l \frac{d^{2} \lambda}{d t^{2}}+l_{1} \frac{d^{2} \alpha}{d t^{2}}\right)=Q_{\lambda}, \quad m l\left(l \frac{d^{2} \mu}{d t^{2}}+l_{1} \frac{d^{2} \beta}{d t^{2}}\right)=Q_{\mu} \\
A \frac{d^{2} \alpha}{d t^{2}}+C \beta \frac{d^{2} \varphi}{d t^{2}}+C \frac{d \varphi}{d t} \frac{d \beta}{d t}+m l_{1}\left(l \frac{d^{2} \lambda}{d t^{2}}+l_{1} \frac{d^{2} \alpha}{d t^{2}}\right)=Q_{\alpha}  \tag{2.6}\\
A \frac{d^{2} \beta}{d t^{2}}-C \frac{d \alpha}{d t} \frac{d \varphi}{d t}+m l_{1}\left(l \frac{d^{2} \mu}{d t^{2}}+l_{1} \frac{d^{2} \beta}{d t^{2}}\right)=Q_{\beta} \\
C \frac{d}{d t} \frac{d \varphi}{d t}=Q_{\varphi}
\end{gather*}
$$

3. The right hand sides of the equations (2.6) represent generalized forces which correspond to selected independent coordinates of the rigid body. These quantities appear as coefficients on perturbing the corresponding generalized coordinates appearing in the expression for the elementary work $\delta W$ done by active forces acting on the body during its arbitrary displacement. We have [4]

$$
\begin{equation*}
\delta W=\mathbf{P} \delta \mathbf{r}_{0}+\mathbf{L}_{0} \delta \boldsymbol{\gamma} \tag{3.1}
\end{equation*}
$$

where $\mathbf{P}$ is the principal vector of all the active forces applied to the body and $\mathrm{L}_{0}$ is their principal moment with respect to the point $O ; \delta \mathrm{r}_{0}$ is a possible displacement of the point $O, \delta y$ is the vector of possible angular displacement of the body.

We shall denote by $P_{\xi}, P_{n}$, and $P_{\zeta}$, the projections of the vector $\mathbf{P}$ on the axes $\xi, \eta$ and $\zeta_{\text {and }}$ by $L_{a}, L_{\beta}$, and $L_{\varphi}$ the sums of the moments of the active forces with respect to the axes $\xi^{\circ}, \eta^{\prime}$ and $\Sigma$ (Fig. 3).

Then, for small angles $\alpha$ and $\beta$ we shall obtain the following expression for $\delta W$

$$
\begin{equation*}
\delta W=P_{\xi} \delta \xi_{0}+P_{n} \delta \eta_{0}+P_{\zeta} \delta \xi_{0}+L_{\alpha} \delta \alpha+L_{\beta} \delta \beta+L_{\varphi} \delta \varphi \tag{3.2}
\end{equation*}
$$

or, using (2.3) and assuming that the angles $\lambda$ and $\mu$ are also small quantities

$$
\begin{equation*}
\delta W=l\left(P_{n}+\lambda P_{\zeta}\right) \delta \lambda+l\left(P_{\zeta} \mu-P_{\xi}\right) \delta \mu+L_{\alpha} \delta \alpha+L_{\beta} \delta \beta+L_{\varphi} \delta \varphi \tag{3.3}
\end{equation*}
$$

From the last expression it follows

$$
\begin{equation*}
Q_{\lambda}=l\left(P_{n}+\lambda P_{\zeta}\right), Q_{\mu}=l\left(P_{\zeta} \mu-P_{\xi}\right), Q_{\alpha}=L_{\alpha}, Q_{\beta}=L_{\beta}, Q_{\varphi}=L_{\varphi} \tag{3.4}
\end{equation*}
$$

The active forces acting on the body are the force of gravity $m g$ and the forces exerted on the body by the enclosed fluid. From this it follows that the generalized force $Q_{\varphi}$ is equal to zero. In fact, it represents the sum of the moments of the force of gravity and of the forces due to the pressure of the fluid with respect to the $z$-axis. However, because of the symmetry, the center of gravity of the body is on the $z$-axis and the vector representing the pressure of the fluid on the body crosses this axis everywhere.

Considering the cosine table (2.1), the formulas (3.3) and the smallness of the angles $\alpha, \beta, \lambda$ and $\mu$, the equations of motion (2.6) can be written in the form

$$
\begin{equation*}
m l\left(l \frac{d^{2} \lambda}{d t^{2}}+l_{1} \frac{d^{2} \alpha}{d t^{2}}\right)=-m g l \lambda+l\left[F_{x} \sin \varphi+F_{y} \cos \varphi-F_{z} \alpha\right]+l F_{2} \lambda \tag{3.5}
\end{equation*}
$$

$$
\begin{gathered}
m l\left(l \frac{d^{2} \mu}{d t^{2}}+l_{1} \frac{d^{2} \beta}{d t^{2}}\right)=-m g l \mu-l\left[F_{x} \cos \varphi-F_{y} \sin \varphi+F_{z} \beta\right]+l F_{z} \mu \\
\left(A+m l_{1}^{2}\right) \frac{d^{2} \alpha}{d t^{2}}+m l_{1} l \frac{d^{2} \lambda}{d t^{2}}+C \frac{d \varphi}{d t} \frac{d \beta}{d t}=-m g l_{1} \alpha+M_{x} \cos \varphi-M_{y} \sin \varphi \\
\left(A+m l_{1}^{2}\right) \frac{d^{2} \beta}{d t^{2}}+m l_{1} l \frac{d^{2} \mu}{d t^{2}}-C \frac{d \varphi}{d t} \frac{d \alpha}{d t}=-m g l_{1} \beta+M_{x} \sin \varphi+M_{y} \cos \varphi \\
C \frac{d}{d t} \frac{d \varphi}{d t}=0
\end{gathered}
$$

In these equations, besides the quantities already known, $g$ is the acceleration due to gravity and

$$
\begin{equation*}
F_{x}=\iint_{\sigma} p^{*} \cos x v d \sigma, \quad F_{y}=\iint_{\sigma} p^{*} \cos y v d \sigma, \quad F_{z}=\iint_{\sigma} p^{*} \cos z v d \sigma \tag{3.6}
\end{equation*}
$$

are the projections on the axes $x, y$, and $z$ of the principal vector of the pressure forces exerted by the flaid on the body.

Similarly

$$
\begin{equation*}
M_{x}=\iint_{\sigma}\left(p^{*} y \cos z v-p^{*} z \cos y v\right) d \sigma, \quad M_{y}=\iint_{\sigma}\left(p^{*} z \cos x v-p^{*} x \cos z v\right) d \sigma \tag{3.7}
\end{equation*}
$$

are the projections on the axes $x$ and $y$ of the principal moment of the forces mentioned earlier, with respect to the point of suspension $O$. The integration in the relations (3.6) and (3.7) is made along the side of the cavity $\sigma ; \cos x v, \cos y v$ and $\cos z v$ are the direction cosines of the external normal $\nu$ to the surface of the cavity; $x, y$ and $z$ are the running coordinates of the element of surface $d \sigma$.

The formulas (3.6) and (3.7) see [5]) can be replaced by the following

$$
\begin{gather*}
F_{x}=\iiint_{\tau} \frac{\partial p^{*}}{\partial x} d \tau, \quad F_{y}=\iiint_{\tau} \frac{\partial p^{*}}{\partial y} d \tau ; \quad F_{z}=\iint_{\tau} \frac{\partial p^{*}}{\partial z} d \tau  \tag{3.8}\\
M_{x}=\iiint_{\tau}\left(y \frac{\partial p^{*}}{\partial z}-z \frac{\partial p^{*}}{\partial y}\right) d \tau, \quad M_{u}=\iint_{\tau}\left(z \frac{\partial p^{*}}{\partial x}-x \frac{\partial p^{*}}{\partial z}\right) d \tau \tag{3.9}
\end{gather*}
$$

Here, the integration is performed over the whole volume of the cavity.
4. In order to determine the pressure $p^{*}$ inside the fluid in terms of the coordinates $x, y$ and $z$ and the time $t$, we shall consider the equations of motion of the fluid with respect to the moving system of coordinates $x y z$ stationary with respect to the rigid body. In those equations we shall consider the projections $u_{x}, u_{y}$, and $u_{z}$ of the relative velocity of any particle of the fluidgand their derivatives with respect to the coordinates, to be small. Disregading, furthermore, the products of small quantities, we get

$$
\begin{equation*}
\frac{\partial u_{x}}{\partial t}=-\frac{1}{\rho} \frac{\partial p^{*}}{\partial x}-w_{x}^{e}-w_{x}^{c}-g \cos \xi x \tag{4.1}
\end{equation*}
$$

The symbol ( $x y z$ ) me ans that the two other formulas are obtained by circular permutations; $w_{x}{ }^{e}, w_{y}{ }^{e}$, and $w_{z}{ }^{e}$ are the projections on the $x, y$ and $z$ axes of the translational acceleration $w_{x}{ }^{c}, w_{y}{ }^{c}$, and $w_{z}{ }^{c}$ of the Coriolis acceleration of a fluid particle, $\cos \zeta x, \cos \zeta y$, and $\cos \zeta z$ are the direction cosines of the axis $\zeta$ in the $x y z$ system of coordinates. The latter terms are equal to the quantities $\alpha \sin \varphi-\beta \cos \varphi$, $\alpha \cos \varphi+\beta \sin \varphi$ and 1 respectively (Fig. 3 and (2.1)), with the accuracy of up to the second order terms of the small angles $\alpha$ and $\beta$. The projections of translational acceleration are given by the formulas

$$
\begin{equation*}
w_{x}^{e}=w_{x}^{0}+\frac{d \omega_{v}}{d t} z-\frac{d \omega_{z}}{d t} y+\omega_{x}\left(x \omega_{x}+y \omega_{y}+z \omega_{z}\right)-\omega^{2} x \quad \text { (xyz) } \tag{4.2}
\end{equation*}
$$

Here $w_{x}{ }^{0}, w_{y}{ }^{0}$, and $w_{z}^{0}$ are the projections on the axes $x, y$ and $z$ of the total acceleration of the point $O$, which is the origin of $x y z$ coordinate system.

These projections can be expressed with accuracy of up to the second order in terms of the derivatives of the angles $\alpha, \beta, \lambda$ and $\mu$ by
$w_{x}^{0}=-l \frac{d^{2} \mu}{d t^{2}} \cos \varphi+l \frac{d^{2} \lambda}{d t^{2}} \sin \varphi, \quad w_{v}{ }^{0}=l \frac{d^{2} \mu}{d t^{2}} \sin \varphi+l \frac{d^{2} \lambda}{d t^{2}} \cos \varphi, \quad w_{z}^{0}=0$
With the same accuracy we obtain by (2.2),
$\omega_{x}=\frac{d \alpha}{d t} \cos \varphi+\frac{d \beta}{d t} \sin \varphi, \quad \omega_{y}=-\frac{d \alpha}{d t} \sin \varphi+\frac{d \beta}{d t} \cos \varphi, \quad \omega_{z}=\frac{d \varphi}{d t}=\omega$
From the last equation of (3.5) it follows that $\omega=$ const, and if we disregard an unimportant constant, then $\varphi=\omega t$. Neglecting furthermore in the expressions (4,2) the terms of the second order in $\omega_{x}$ and $\omega_{y}$ (which are of the order of $d \alpha / d t$ and $d f . d t$ ), and also the terms which include the derivative of $\omega_{z}$ with respect to time, we get

$$
\begin{gather*}
w_{x}^{e}=-l \frac{d^{2} \mu}{d t^{2}} \cos \varphi+l \frac{d^{2} \lambda}{d t^{2}} \sin \varphi+\frac{d^{2} \omega_{y}}{d t} z+\omega \omega_{x} z-\omega^{2} x \\
w_{y}{ }^{8}=l \frac{d^{2} \mu}{d t^{2}} \sin \varphi+l \frac{d^{2} \lambda}{d t^{2}} \cos \varphi-\frac{d \omega_{x}}{d t} z+\omega \omega_{y} z-\omega^{2} y  \tag{4.5}\\
w_{z}^{e}=\frac{d \omega_{x}}{d t} y-\frac{d \omega_{y}}{d t} x+\omega \omega_{x} x+\omega \omega_{u} y
\end{gather*}
$$

The projections of the Coriolis acceleration on the $x, y$, and $z$ axes are equal [4] to $w_{x}^{c}=2\left(\omega_{y} u_{z}-\omega_{z} u_{y}\right), \quad w_{v}{ }^{c}=2\left(\omega_{2} u_{x}-\omega_{x} u_{z}\right), \quad w_{x}^{c}=2\left(\omega_{x} u_{y}-\omega_{y} u_{x}\right)$

Here also, one should neglect the terms in which $\omega_{x}$ and $\omega_{y}$, appear as multipliers and write $\omega_{z}=\omega$.

Taking the formulas (4.2) and (4.6) into account, the equations (4.1) take the form

$$
\begin{gather*}
\frac{\partial u_{x}}{\partial t}-2 \omega u_{y}=-\frac{1}{\rho} \frac{\partial p_{1}}{\partial x}-2 z \frac{d \omega_{y}}{d t} \\
\frac{\partial u_{y}}{\partial t}+2 \omega u_{x}=-\frac{1}{\rho} \frac{\partial p_{1}}{\partial y}+2 z \frac{d \omega_{x}}{d t}, \frac{\partial u_{x}}{\partial t}=-\frac{1}{\rho} \frac{\partial p_{1}}{\partial z} \tag{4.7}
\end{gather*}
$$

Where

$$
\begin{align*}
& p_{1}=p^{*}+\rho\left[\omega z\left(x \omega_{x}+y \omega_{y}\right)-1 / 2 \omega^{2}\left(x^{2}+y^{2}\right)\right]+\rho g[x(\alpha \sin \varphi-\beta \cos \varphi)+  \tag{4.8}\\
& +y(\alpha \cos \varphi+\beta \sin \varphi)+z]+\rho\left(x w_{x}^{0}+y w_{y}{ }^{0}+z w_{z}^{0}\right)+\rho z\left(y \frac{d \omega_{x}}{d t}-x \frac{d \omega_{y}}{d t}\right)
\end{align*}
$$

and $\omega_{x}, \omega_{y}, w_{x}{ }^{0}, w_{v}{ }^{0}$, and $w_{z}{ }^{0}$ are determined by the equations (4.3) and (4.4).
The functions $u_{x}, u_{y}$, and $u_{z}$ must, farthermore, satisfy the condition of incompressibility

$$
\begin{equation*}
\frac{\partial u_{x}}{\partial x}+\frac{\partial u_{y}}{\partial y}+\frac{\partial u_{z}}{\partial z}=0 \tag{4.9}
\end{equation*}
$$

and the boundary condition

$$
\begin{equation*}
u_{x} \cos x v+u_{y} \cos y v+u_{z} \cos z v=0 \tag{4.10}
\end{equation*}
$$

The latter equation means that the projection of the relative velocity of a fluid particle in contact with the boundary of the cavity, onto the axis normal to this bonndary, is equal to zero.
5. Following the work of Sobolev [3], we search for $u_{x}, u_{y}$ and $u_{z}$ in the form $u_{x}=U_{x}(t)\left(z+l_{2}\right), \quad u_{y}=U_{y}(t)\left(z+l_{2}\right), \quad u_{z}=-\frac{c^{2}}{a^{2}}\left[x U_{x}(t)+y U_{y}(t)\right]$

Here $U_{x}(t)$ and $U_{y}(t)$ are determinable functions of time*; $l_{2}$ is the distance between the center of the ellipsoidal cavity and the origin of the $x y z$ coordinate system (Fig. 1).

We denote the major and the minor semi-axe of the ellipsoidal cavity by $a$ and $c$ respectively. Then, its equation will take the form

$$
\begin{equation*}
\frac{x^{2}+y^{2}}{a^{2}}+\frac{\left(z+l_{2}\right)^{2}}{c^{2}}=1 \tag{5.2}
\end{equation*}
$$

and the direction cosines of the normal $\nu$ in the $x y x$ coordinate system will be

[^0]\[

$$
\begin{gather*}
\cos x v=\frac{x c^{2}}{D}, \quad \cos y v=\frac{y c^{2}}{D}, \cos z v=\frac{a^{2}\left(z+l_{2}\right)}{D}  \tag{5.3}\\
D=\sqrt{c^{4}\left(x^{2}+y^{2}\right)+a^{4}\left(z+l_{2}\right)^{2}} \tag{5.4}
\end{gather*}
$$
\]

By direct substitution of the relations (5.1) and (5.3) into the boundary condition (4.10) and using the condition of incompressibility (4.9) we can see that the latter is identically satisfied.

Let us now consider the system (4.7). Multiplying its first equation by $\cos x \nu$, the second by cos $y \nu$ and the third by $\cos x \nu$ and adding, we obtain by utilising also (5.1) and (5.3).

$$
\begin{gather*}
\frac{1}{\mathrm{P}} \frac{\partial p_{1}}{\partial n}=\frac{2 c^{2}}{D} z\left\{x\left[\omega U_{y}(t)-\frac{d \omega_{y}}{d i}\right]+y\left[-\omega U_{x}(t)+\frac{d \omega_{x}}{d t}\right]\right\}+  \tag{5.5}\\
+\frac{2 c^{2} l_{2} \omega}{D}\left[x U_{y}(t)-y U_{x}(t)\right]
\end{gather*}
$$

The last relation becomes an identity if we write

$$
\begin{equation*}
p_{1}=\rho z\left[x P^{*}(t)+y Q^{*}(t)\right]+\rho\left[x P_{1}^{*}(t)+y Q_{1}^{*}(t)\right] \tag{5.6}
\end{equation*}
$$

and choose the functions $P^{*}(t), P_{1}{ }^{*}(t), Q^{*}(t)$ and $Q_{1}{ }^{*}(t)$ as follows:
$p^{*}(t)=\frac{2 c^{2}}{a^{2}+c^{2}}\left[\omega U_{y}(t)-\frac{d \omega_{y}}{d t}\right], \quad p_{1}^{*}(t)=\frac{2 l_{2}}{a^{2}+c^{2}}\left[a^{2} \frac{d \omega_{y}}{d t}+\omega U_{y}(t) c^{2}\right]$
$Q^{*}(t)=\frac{2 c^{2}}{a^{2}+c^{2}}\left[-\omega U_{x}(t)+\frac{d \omega_{x}}{d t}\right], \quad Q_{1}^{*}(t)=-\frac{2 l_{2}}{a^{2}+c^{2}}\left[a^{2} \frac{d \omega_{x}}{d t}+\omega U_{x}(t) c^{2}\right]$
Let us now multiply the second equation of the system (4.7) by $i=\sqrt{-1}$ and add it to the first. Then by using the equalities (5.1), (5.6), (5.7) and simplyfing by the factor $z+l_{2}$ we obtain the relation necessary for the determination of the functions $U_{x}(t)$ and $U_{y}(t)$, i.e.

$$
\left[\left(a^{2}+c^{2}\right) \frac{d}{d t}+2 \omega a^{2} i\right]\left[U_{x}(t)+i U_{y}(t)\right]=2 a^{2} i\left(\begin{array}{c}
d \omega_{x}  \tag{5.8}\\
d t
\end{array}+i \frac{d \omega_{y}}{d t}\right)
$$

6. Now, let us investigate the system (3.5). Let us multiply its second and fourth equations by $i=\sqrt{-1}$ and add the first to the second and the third to the fourth. Then, let us also introduce the complex functions of the real variable $t$

$$
\begin{equation*}
\zeta^{*}=\alpha+i \beta, \quad z^{*}=\lambda+i \mu \tag{6.1}
\end{equation*}
$$

Taking into congideration the third equality of (4.4) we have

$$
\begin{gather*}
\left(A+m l_{1}^{2}\right) \frac{d^{2} \zeta^{*}}{d t^{2}}-i C \omega \frac{d^{*}}{d t}+m l_{1} l \frac{d^{2} z^{*}}{d t^{2}}+m g l_{1} \zeta^{*}=\left(M_{x}+i M_{y}\right) e^{i \omega t} \\
m\left(l \frac{d^{2} z^{*}}{d t^{2}}+l_{1} \frac{d^{2} \zeta^{*}}{d t^{2}}\right)+\left(m g-F_{z}\right) z^{*}=-i\left(F_{x}+i F_{y}\right) e^{i \omega t}-F_{z} \zeta^{*} \tag{6.2}
\end{gather*}
$$

To calculate the expressions $F_{x}+i F_{y}, F_{z}$, and $M_{x}+i M_{y}$, which appear in the right hand sides of the system (6.2), we substitute into the formulas (3.8) and (3.9) the expressions for the pressure $p^{*}$ found by eliminating the variable $p_{1}$ between the relations (4.8) and (5.6). Also we note, that

$$
\begin{align*}
& \iiint_{\tau} \rho x z d \tau=\iiint_{\tau} \rho x y d \tau=\iiint_{\tau} \rho z y d \tau=0, \quad \iiint_{\tau} \rho\left(x^{2}-y^{2}\right) d \tau=0 \\
& \iiint_{\tau} \rho x d \tau=\iiint_{\tau} \rho y d \tau=0, \quad \iiint_{\tau} \rho d \tau=m_{1}, \quad \iiint_{\tau} \rho z d \tau=-m_{1} l_{2}  \tag{6.3}\\
& \iiint_{\tau} \rho\left(z^{2}-y^{2}\right) d \tau=\iiint_{\tau} \rho\left(z^{2}-x^{2}\right) d \tau=\frac{4}{15} \pi \rho a^{2} c\left(c^{2}-a^{2}\right)+m_{1} l_{2}^{2}
\end{align*}
$$

Here, in addition to the symbols already seen earlier, $\rho$ and $m_{1}$ represent respectively the specific density and the mass of the fluid which fills the ellipsoidal cavity of the rigid body.

If, furthermore, we notice that according to the formula (4.3), (4.4) and (6.1)

$$
\begin{equation*}
\omega_{x}+i \omega_{y}=\frac{d \zeta^{*}}{d t} e^{-i \omega t}, \quad w_{x}^{0}+i w_{y}^{0}=i l \frac{d^{2} z^{*}}{d t^{2}} e^{-i \omega t} \tag{6.4}
\end{equation*}
$$

then, after a few rather simple calculations and the use of the equalities (5.7), (5.8) we get the required expressions of $F_{x}+i F_{\nu}$ and $F_{z}$

$$
\begin{equation*}
F_{x}+i F_{y}=-m_{1} i\left[l \frac{d^{2} z^{*}}{d t^{2}}+l_{2} \frac{d^{2} \zeta^{*}}{d t^{2}}+g \zeta^{*}\right] e^{-i \omega t}, \quad F_{z}=-m_{1} g \tag{6.5}
\end{equation*}
$$

and also the differential equation for $M_{x}+i M_{y}$, which is

$$
\begin{gather*}
M_{x}+i M_{y}=k \frac{-\left(c^{2}-a^{2}\right) d^{3} \zeta^{*} / d t^{3}+2 \omega i c^{2} d^{2} \zeta^{*} / d t^{2}+\left(a^{2}+c^{2}\right) \omega^{2} d \zeta^{*} / d t}{\left(a^{2}+c^{2}\right) d / d t+2 \omega a^{2} i} e^{-i \omega t}- \\
\quad-\left[m_{1} l_{2}{ }^{d^{2} \zeta^{*}} \frac{d t^{2}}{d t_{1}} m_{1} g l_{2} \zeta^{*}+m_{1} l_{2} l \frac{d^{2} z^{*}}{d t^{2}}+i \omega k \frac{d \zeta^{*}}{d t}\right] e^{-i \omega t} \tag{6.6}
\end{gather*}
$$

where

$$
\begin{equation*}
k=4 / 15 \pi \rho a^{2} c\left(c^{2}-a^{2}\right) \tag{6.7}
\end{equation*}
$$

7. Eliminating $F_{x}+i F_{y}, F_{z}$ and $M_{x}+i M_{y}$ from (6.2) by using the relation (6.5) and (6.6) we obtain the following system of equations with respect to the complex functions of time $\zeta^{*}$ and $z^{*}$

$$
\begin{gather*}
\left(A^{*}+k \eta\right) \frac{d^{3} \zeta^{*}}{d t^{3}}-i \omega\left[C+\left(A^{*}+k\right) \eta\right] \frac{d^{2} \zeta^{*}}{d t^{2}}+\left[K-\omega^{2}(C \eta-k \eta+k)\right] \frac{d \zeta^{*}}{d t}- \\
-i \omega \eta K \zeta^{*}+\left(m l_{1}+m_{1} l_{2}\right) l \frac{d^{3} z^{*}}{d t^{3}}-i \omega \eta l\left(m l_{1}+m_{1} l_{2}\right) \frac{d^{3 z^{*}}}{d t^{2}}=0  \tag{7.1}\\
l \frac{d^{2} z^{*}}{d t^{2}}+z_{c} \frac{d^{2} \zeta^{*}}{d t^{2}}+g z^{*}=0
\end{gather*}
$$

Where

$$
\begin{align*}
A^{*} & =A+m l_{1}{ }^{2}+m_{1} l_{2}{ }^{2}, \quad \eta=\frac{c^{2}-a^{2}}{c^{2}+a^{2}}, \quad z_{c}=\frac{m l_{1}+m_{1} l_{2}}{m+m_{1}} \tag{7.2}
\end{align*}
$$

We assume that the solution of (7.1) has the form

$$
\begin{equation*}
\zeta^{*}=\zeta^{0} e^{i \lambda, t}, \quad z^{*}=z^{0} e^{i \lambda, t} \tag{7.3}
\end{equation*}
$$

At the same time, the characteristic equation reduces to the following:

$$
\begin{align*}
f(\lambda ; \omega) & =\left(A^{0}+k \eta\right) \lambda^{5}-\omega\left[C+\left(A^{0}+k\right) \eta\right] \lambda^{4}+\left[-K-\left(A^{*}+k \eta\right) \frac{g}{l}+\right. \\
& \left.+C \omega^{2} \eta+k \omega^{2}(1-\eta)\right] \lambda^{3}+\left[K \eta+C \frac{g}{l}+\frac{g}{l}\left(A^{*}+k\right) \eta\right] \omega \lambda^{2}+(7.4)  \tag{7.4}\\
& +\frac{g}{l}\left[K-k \omega^{2}(1-\eta)-C \omega^{2} \eta\right] \lambda-K \omega \eta \frac{g}{l}=0
\end{align*}
$$

Where

$$
\begin{equation*}
A^{0}=A^{*}-z_{c}\left(m l_{1}+m_{1} l_{2}\right) \tag{7.5}
\end{equation*}
$$

The equations (7.1) together with the equality $\omega_{z}=\omega=$ const describe the motion of the body in question. Hence the investigation of the stability of its motion reduces to the study of behaviour of the functions (7.3). Obviously, the condition that the roots of the characteristic equation (7.4) are real, is the criterion of the stability of motion. Below, we shall try to determine this condition.
8. First, we shall consider a few particular cases. a) The length of the string $l \rightarrow 0$. Then, the characteristic equation (7.4) takes the form
$\left(A^{*}+k \eta\right) \lambda^{3}-\left[C+\left(A^{*}+k\right) \eta\right] \omega \lambda^{2}-\left[K-k \omega^{2}(1-\eta)-C \omega^{\circ} \eta\right] \lambda+K \omega \eta=0(8.1)$
Note, that even as far as the notations are concerned, the equation (8.1) coincides with the equation obtained in papers [1] and [3] in the investigation of the stability of a gyroscope with an ellipsoidal cavity completely filled with an ideal fluid.* b) The cavity has a spherical form. Then $a=c$ and, according to equation (6.7) and the third relation of (7.2), $k=0, \eta=0$. The characteristic equation (7.4) takes the form
$\lambda\left\{\lambda^{4}-\frac{C}{A^{0}} \omega \lambda^{3}-\frac{g}{l}\left[1+\frac{\left(m l_{1}+m_{1} l_{2}\right)\left(l+z_{c}\right)}{A^{0}}\right] \lambda^{2}+\frac{C}{A^{0}} \omega \frac{g}{l} \lambda+\frac{g^{2}}{l} \frac{m l_{1}+m_{1} l_{2}}{A^{0}}\right\}=0$
One of the roots of the equation (8.2) is equal to zero, and the remaining four roots, can easily be shown to be located on the intervals

$$
(-\infty,-\sqrt{g / l}),(-\sqrt{g / l}, 0)(0, \sqrt{g / l}),(\sqrt{g / l}, \infty)
$$

Thus, the motion of the body possessing a spherical cavity completely filled with an ideal fluid, is always stable. It has the same character, as the motion of a dense rigid body undergoing a rotation, investigated in [2].

## 9. Let us now return to the study of the stability of the motion of a rigid body with

[^1] investigated in [1] and [3].
an ellipsoidal cavity completely filled with an ideal incompressible fluid. For that purpose we consider the characteristic equation (7.4) and attempt to determine the conditions for its roots to be real (i.e. the conditions of stability of motion).

A series of criteria are available for determining real roots of algebraic equations of the $n$-th order (See for instance [6]). However, in the present case (because of the complexity of the coefficients of the equations (7.4), these criteria are very cumbersome and consequently of little use for the study of stability conditions for the body under investigation, when the parameters entering the equation (7.1) of its motion, are subject to an arbitrary variation. Below is given a graphical-analytical method of investigation of the roots of the equation (7.4). It enables us to determine easily the zones of stable and unstable motion of the body in question for arbitrary values of its angular velocity $\omega$.

In agreement with the rule of Descartes [7], the number of positive roots of the equation (7.4), for any values of the parameter $\omega>0$, cannot be greater than three* and the number of negative roots cannot be greater than two. It will be shown later that the equation (7.4) for $\omega>0$ has always two negative roots.

In fact, taking into consideration the equalities (7.2) and (7.5) we have

$$
\begin{equation*}
f(-\sqrt{g / l}, \omega)=\frac{g^{2}}{l^{2}}\left(A^{*}-A^{\circ}\right)(\omega \eta \sqrt{g / l})>0 \tag{9.1}
\end{equation*}
$$

However, $f(-\infty ; \omega)<0$ and $f(0 ; \omega)<0$. Consequently, in each of the intervals $-\infty<\lambda<-\sqrt{g}-l$ and $-\sqrt{g / l}<\lambda<0$, there is one (negative) root of the equation (7.4).

To determine the character of the three ramaining roots of the equation (7.4) let us collect the terms of its left hand side in decreasing powers of $\omega$ i.e.

$$
\begin{equation*}
p(\lambda) \omega^{2}-q(\lambda) \omega+r(\lambda)=0 \tag{9.2}
\end{equation*}
$$

Where

$$
\begin{align*}
p(\lambda)= & \lambda\left(\lambda^{2}-\frac{g}{l}\right) \varepsilon \eta^{2}, q(\lambda)=\eta\left[R(\lambda)+\lambda^{2}\left(\lambda^{2}-\frac{g}{l}\right)(\varepsilon-x)\right], r(\lambda)=\lambda R(\lambda)(9.3) \\
& R(\lambda)=\left(A^{0}+k \eta\right) \lambda^{4}-\left[K+\frac{g}{l}\left(A^{*}+k \eta\right)\right] \lambda^{2}+\frac{g}{l} K \\
\varepsilon= & \frac{C \eta+k(1-\eta)}{\eta^{2}}>0, \quad x=\frac{k(1-\eta)\left(1-\eta^{2}\right)}{\eta^{2}}>0, \quad(\varepsilon-x)>0 \tag{9.4}
\end{align*}
$$

Curves representing the polynomials $p(\lambda), q(\lambda)$, and $r(\lambda)$ for $\lambda \geqslant 0$ are shown in Fig. 4. The analysis of the curves obtained shows that for $\lambda \geqslant 0$, the polynomial $r(\lambda)$ becomes equal to zero at $\lambda=0$ and at two other values of $\lambda$ denoted by $\lambda_{1}$ and $\lambda_{2}$. Similarly the polynomial $q(\lambda)$ is equal to zero only when $\lambda=\lambda_{1}{ }^{0}$, and $\lambda=\lambda_{2}{ }^{0}$. Finally, the polynomial $p(\lambda)$ is equal to zero either when $\lambda=0$, or $\lambda=\sqrt{g} / \sqrt{l}$.

[^2]

FIG. 4

It is possible to solve the equation (9.2) for $\omega$ and construct the curve representing the function

$$
\begin{equation*}
\omega(\lambda)= \tag{9.5}
\end{equation*}
$$

$=\frac{q(\lambda) \pm \sqrt{q^{2}(\lambda)-4 p(\lambda) r(\lambda)}}{2 p(\lambda)}$
From this curve, it is possible to establish for the given value of the angular velocity $\omega$, whether all the roots of the equation (7.4) with respect to $\lambda$ are real and, consequently, whether the motion of a rigid body with a fluid filled cavity is stable, or not.

Within the range of values of $\lambda$ for which the discriminant

$$
\begin{equation*}
\Delta(\lambda)=4 p(\lambda) r(\lambda)-q^{2}(\lambda) \tag{9.6}
\end{equation*}
$$

is positive, the valnes of $\omega$, according to the formula (9.5) are complex conjugate quantities. This will occur in the intervals $\left(\lambda_{1}{ }^{*}, \lambda_{1}{ }^{* *}\right)$ and $\left(\lambda_{2}{ }^{*}, \lambda_{2}{ }^{* *}\right)$, where $\lambda_{1}{ }^{*}, \lambda_{2}{ }^{*}$, and $\lambda_{1}{ }^{* *}, \lambda_{2}{ }^{* *}$ are the positive roots of the polynomial $\Delta(\lambda)$; these roots (as well as the negative roots possessing the same moduli) exist for any values of the parameters entering the expression (9.6). This can be proven by expressing the discriminant $\Delta(\lambda)$ in the form of a product

$$
\begin{gather*}
\Delta(\lambda)=-\eta^{2}\left[R(\lambda)-\lambda^{2}\left(\lambda^{2}-\frac{g}{l}\right)(\sqrt{\varepsilon}+\sqrt{x})^{2}\right] \times \\
\times\left[R(\lambda)-\lambda^{2}\left(\lambda^{2}-\frac{g}{l}\right)(\sqrt{\varepsilon}-\sqrt{x})^{2}\right] \tag{9.7}
\end{gather*}
$$

and taking into account the form of the polynomial $R(\lambda)$ according to the first formula (9.4).
According to the formula (9.5) both values of $\omega$ become equal for the values of $\lambda$, which coincide with one of the roots of the discriminant $\Delta(\lambda)$. Let us denote them by $\omega_{1}{ }^{*}, \omega_{2}{ }^{*}, \omega_{1}{ }^{* *}$ and $\omega_{2}{ }^{* *}$. We find that they correspond to the cusps on the curve $\omega=\omega(\lambda)$. This curve is shown in Fig. 5. When constructing it, one must take into consideration the distribution of the zeros of the polynomials $p(\lambda), q(\lambda)$ and $r(\lambda)$ determined by the formulas (9.3) and by the Fig. 4. In the intervals ( $\omega_{1}{ }^{*}, \omega_{1}{ }^{* *}$ ) and ( $\omega_{2}{ }^{*}$, $\omega_{2}{ }^{* *}$ ) (see Fig. 5), the fifth order equation (7.4) in $\lambda$ has only three real roots (two of which are negative) for each calue of $\omega$. The remaining two roots are complex.

Following the exposition made above, we come to the following conclusion : the motion of a rigid body with en ellipsoidal cavity completely filled with an ideal incompressible
fluid is unstable for values of the angular velocity varying between the limits $\omega_{1}{ }^{*}<\omega<\omega_{1}{ }^{* *}$ and $\omega_{2}{ }^{*}<\omega<\omega_{2}{ }^{* *}$. Outside these intervals the motion of the body in question is stable.
10. The critical values $\omega_{1}{ }^{*}, \omega_{1}{ }^{* *}, \omega_{2}{ }^{*}$, and $\omega_{2}{ }^{* *}$ of the angular velocity are determined by the expression (9.5) in which we put

$$
q^{2}(\lambda)-4 p(\lambda) r(\lambda)=-\Delta(\lambda)=0
$$

This results in the following simple formulas

$$
\begin{array}{rlrl}
\omega_{1}^{*} & =\frac{q\left(\lambda_{1}{ }^{*}\right)}{2 p\left(\lambda_{1}^{*}\right)}=\frac{\varepsilon-\sqrt{\varepsilon x}}{\varepsilon \eta} \lambda_{1}^{*}, & \omega_{2}^{*}=\frac{q\left(\lambda_{2}^{*}\right)}{2 p\left(\lambda_{2}^{*}\right)}=\frac{\varepsilon-\sqrt{\varepsilon x}}{\varepsilon \eta} \lambda_{2}^{*} \\
\omega_{1}^{* *}=\frac{q\left(\lambda_{1}^{* *}\right)}{2 p\left(\lambda_{1}^{* *}\right)}=\frac{\varepsilon+\sqrt{\varepsilon x}}{\varepsilon \eta} \lambda_{1}^{* *}, & \omega_{2}^{* *}=\frac{q\left(\lambda_{2}^{* *}\right)}{2 p\left(\lambda_{2}^{* *}\right)}=\frac{\varepsilon+\sqrt{\varepsilon \chi}}{\varepsilon \eta} \lambda_{2}^{* *} \tag{10.1}
\end{array}
$$

in which, as mentioned before $\lambda_{1}{ }^{*}, \lambda_{1}{ }^{* *}, \lambda_{2}{ }^{*}$ and $\lambda_{2}{ }^{* *}$ are respective positive roots of the discriminant $\Delta(\lambda)$.
11. As an example, we shall determine the critical values of the angular velocity $\omega$ of the rotation of a rigid body, the parameters of which are given below :

$$
\begin{aligned}
A=15.84, C=3.74\left[\mathrm{gram} \mathrm{sec}^{2}\right], m & =0.8145, m_{1}=0.0334\left[\mathrm{gram} \mathrm{sec}^{2} \mathrm{~cm}^{-1}\right] \\
a & =1.5, c=4.35, l_{1}=6.6, l_{2}=5.5, l=50[\mathrm{~cm}]
\end{aligned}
$$

In this case, by the formulas (10.1) $\omega_{1}{ }^{*}=48, \omega_{1}{ }^{* *}=53, \omega_{2}{ }^{*}=264$, and $\omega_{2}{ }^{* *}=309$ [rev/min]. Thus, for values of $\omega 48<\omega<53$, and $264<\omega<309 \mathrm{rev} / \mathrm{min}$, the motion of the body in question is anstable. For all values of $\omega$ outside these intervals it is stable.

The critical values of the angular velocity were also calculated for the same parametric values by using the criterion givenin [6]. They were found to be equal to those determined above, their determination however, required a mach greater amount of work.
'12. In order to check the criterion of stability obtained above for the motion of a rigid body with a flaid-filled cavity, experimental investigations described in [8], were made at the physico-technical laboratory of the Institute of Mechanics of the Academy of Sciences of the Ukrainian SSR by E.V. Virt, and A.P. Polyvianna ander the sapersions of Dr. Malashenko.

The model which consisted of a hollow body which had cylindrical insert inside it was fixed to the axle of a vertical motor by means of a thin string or a capron thread.

The insert consisted of two separate parts which could be hermetically sealed to produce an ellipsoidal cavity (see Fig. 7 in [8]). In the top part of the insert there was an opening through which the cavity could be filled with fluid (in the experiments described, ethanol was used).


FIG. 5

The angular velocity of the model was varied between the limits of 100 and 3000 rev/min. The stabilization of the angular velocity was accomplished by means of a precision sound generator and a thyrathron.

The stability (or instability) of the motion of the model at various values of the angular velocity was determined by tapping the model slightly and observing the resulting behavior.

Experimental results with the model considered in section 11 have shown that the motion of the model was atable over the interval of $100-200 \mathrm{rev} / \mathrm{min}$. In the $220-450 \mathrm{rev} /$ $\min$ range it lost the stability but regained if above $450 \mathrm{rev} / \mathrm{min}$.

The experimental determination of the limits of the first region of critical values of the angular velocity (see section 11 of this work) was not attempted. (The experimental set up could not be used for angular velocities less than $100 \mathrm{rev} / \mathrm{min}$ ).

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[^0]:    * Also ace the note at the beginning of the work [1].

[^1]:    * With the reservation that the sign of the momentum $K$ is opposite to that used in the case

[^2]:    * In case of $\omega=0$, four roots of the equation (7.4) are real, and the fifth is equal to zero.

